

Resurgence and instantons in the $O(N)$ sigma model

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- Perturbation theory in quantum field theory
- Borel resummation
- The $O(N)$ nonlinear σ -model in $1 + 1$ dimension
- Wiener-Hopf solution
- The full trans-series and median resummation

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Successes and problems of perturbation theory in quantum field theory

Anomalous magnetic moment in quantum electrodynamics

g -factor: magnetic moment of elementary $1/2$ spin fermions
(in Bohr-units)

Dirac equation:

$$g = 2$$

quantum field theoretic corrections:

$$a = \frac{g-2}{2}$$

electron:

$$a_e = \sum_{n=1}^5 c_n^{(e)} \left(\frac{\alpha}{\pi}\right)^n$$

$$\alpha \sim \frac{1}{137} \quad \text{fine structure constant}$$

$$c_1^{(e)} = \frac{1}{2} \quad c_2^{(e)} = \frac{197}{144} + \frac{\pi^2}{12} - \frac{\pi^2 \ln 2}{2} + \frac{3}{4}\zeta[3]$$

$c_3^{(e)}$: exactly known (generalized zeta-function)

$c_{4,5}^{(e)}$: known numerically

$$a_e^{\text{theor}} = 0.001\,159\,652\,181\,64(76)$$

T.Aoyama, T.Kinoshita, M.Nio '19

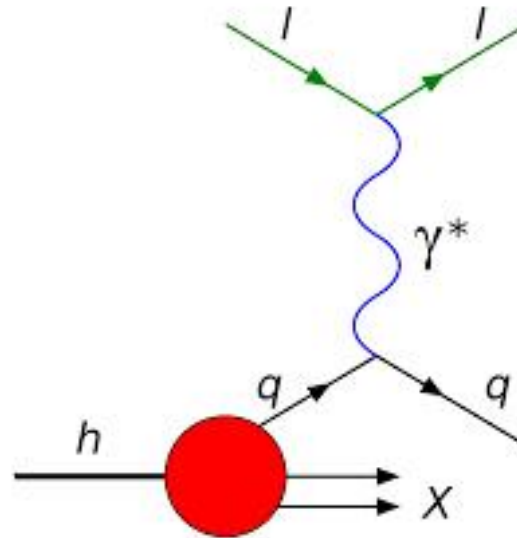
$$a_e^{\text{exp}} = 0.001\,159\,652\,180\,73(28)$$

agree to 10 digits!

muon: “only” 8-digit agreement

3.7σ deviation [new physics?]

Deep inelastic scattering (QCD)



running coupling:

$$\alpha(E^2) \approx \frac{1}{\beta_o \ln\left(\frac{E^2}{\Lambda^2}\right)}$$

QCD Lambda – parameter

first 1-2 (3) terms: $E : 5 - 80\text{GeV}$ $\alpha(E^2) \sim 0.15$

Hadronic τ -decay

hadronic branching ratio:

$$R_\tau = A_{\text{EW}} (|V_{ud}|^2 + |V_{us}|^2) \{1 + \delta_o\}$$

QCD corrections:

$$\delta_o = \frac{1}{2\pi i} \oint_{|s|=m_\tau^2} \frac{ds}{s} \left(1 - \frac{s}{m_\tau^2}\right)^3 \left(1 + \frac{s}{m_\tau}\right) D(s)$$

Adler function:

$$D(s) = \sum_{n=1}^5 c_n^{(\text{A})} \left(\frac{\alpha(s)}{\pi}\right)^n$$

4 (5) loop order:

$$c_1^{(A)} = 1 \quad c_2^{(A)} = 1.640 \quad c_3^{(A)} = 6.371 \quad c_4^{(A)} = 49.076$$

$$c_5^{(A)} = 277 \pm 51$$

P.A.Baikov, K.G.Chetyrkin, J.H.Kühn '08

Practical problem: $\alpha(s) \sim 0.3$ we know too few terms

Issue of principle!

The perturbative series is asymptotic

$n!$ behaviour

typical perturbative series:

$$f = \sum_{n=0}^{\infty} c_n \alpha^n$$

large n asymptotics:

$$c_n \sim n!$$

- number of Feynman diagrams grows like $n!$
- UV and IR “renormalons” (individual diagramm-series) $\sim n!$

How can we obtain the exact value of the physical quantity knowing the perturbative series f ?

Borel resummation

Borel-transform

$$B(t) = \sum_{n=0}^{\infty} \frac{c_n}{n!} t^n \quad \text{convergent}$$

formal Borel summation:

$$\mathcal{S}[f] = \int_0^{\infty} dt e^{-t/\alpha} B(t)$$

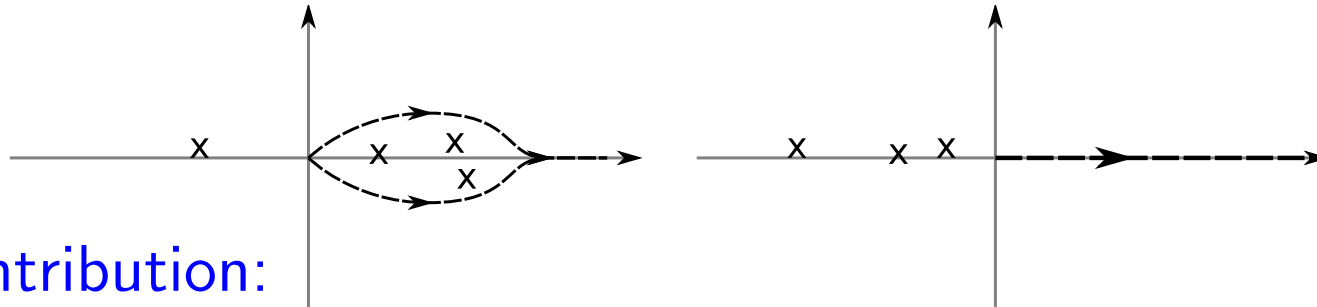
$B(t)$ analytic continuation: singularities on the real axis:

- $t > 0$ IR renormalons and instantons
- $t < 0$ UV renormalons

Lateral Borel-resummation:

$\mathcal{S}_{\pm}[f]$: integration contour slightly above (under) the real axis

- $\mathcal{S}_+[f]$ and $\mathcal{S}_-[f]$ reproduce the original asymptotic series
- **But:** $\mathcal{S}_+[f] \neq \mathcal{S}_-[f]$ **NP ambiguity**



pole contribution:

$$B(t) \sim \frac{b}{t-S}$$

NP ambiguity:

$$\mathcal{S}_+ - \mathcal{S}_- = 2\pi i b e^{-S/\alpha}$$

ambiguous imaginary part: $\text{NP} \sim e^{-S/\alpha}$

‘Naive’ solution:

$$f_{\text{phys}} = \frac{1}{2} (\mathcal{S}_+[f] + \mathcal{S}_-[f]) = \text{Re } \mathcal{S}_+[f]$$

$O(N)$ model in $1 + 1$ dimension

$O(N)$ nonlinear σ -model:

$$\mathcal{L} = \frac{1}{2\alpha_o} \partial_\mu \vec{S} \partial^\mu \vec{S} \quad \vec{S}^2 = 1 \quad \vec{S} = (S_1, \dots, S_N)$$

- renormalizability ($\alpha_o \rightarrow \alpha$) running coupling
- asymptotic freedom
- dynamical mass-generation (m)
- instantons ($N = 3$)
- integrability: exact scattering matrix known

A.B.Zamolodchikov, Al.B.Zamolodchikov '77

We studied the energy $\epsilon(\rho)$ of the finite ρ density state: ($N = 4$)

$$\epsilon(\rho) = \frac{\pi}{2} \rho^2 \alpha f(\alpha)$$

$$f(\alpha) = \sum_{n=0}^{\infty} \chi_n \alpha^n$$

$$\chi_0 = 1 \qquad \chi_1 = \frac{1}{2} \qquad \chi_2 = \frac{1}{4}$$

where

$$\frac{1}{\alpha} - \frac{1}{2} \ln \alpha = \ln \frac{\pi \rho}{\Lambda}$$

TBA equation

$$\chi(\theta) - \int_{-B}^B K(\theta - \theta') \chi(\theta') d\theta' = m \cosh \theta \quad |\theta| \leq B$$

P.Hasenfratz, M. Maggiore, F. Niedermayer '90

- $K(\theta)$: the logarithmic derivative of the S-matrix
- $\chi(\theta)$: particle density in rapidity space
- B : maximal rapidity of the final density state

$$\rho = \frac{1}{2\pi} \int_{-B}^B \chi(\theta) d\theta \quad \epsilon(\rho) = \frac{m}{2\pi} \int_{-B}^B \cosh \theta \chi(\theta) d\theta$$

Volin's trick

D.Volin '10

in the $B \rightarrow \infty$ limit:

$$\rho = A \frac{\sqrt{B}}{\pi} \hat{\rho} \quad \epsilon = m A e^B \frac{1}{4\sqrt{2\pi}} \hat{\epsilon} \quad A = m e^B \frac{\sqrt{\pi}}{2\sqrt{2}}$$

$$\hat{\rho} = \sum_{n=0}^{\infty} r_n B^{-n} \quad \hat{\epsilon} = 1 + \sum_{n=0}^{\infty} e_n B^{-1-n}$$

- numerical coefficients can be calculated recursively, completely algebraically
- $1/B$ expansion can be rewritten as α -expansion:

$$\frac{1}{\alpha} + \frac{1}{2} - B - \frac{1}{2} \ln B\alpha = \ln \hat{\rho}$$

The first 44 perturbative coefficients analytically

Mathematica: first 44 terms

$$\begin{aligned}\chi_0 &= 1 & \chi_1 &= \frac{1}{2} & \chi_2 &= \frac{1}{4} \\ \chi_3 &= \frac{5}{16} - \frac{3}{32}\zeta[3] & \chi_4 &= \frac{53}{96} - \frac{9}{64}\zeta[3]\end{aligned}$$

$\vdots$ $\vdots$

$$\begin{aligned}\chi_8 &= \frac{30150661}{645120} - \frac{134905}{8192}\zeta[3] + \frac{19719}{8192}\zeta[3]^2 \\ &\quad - \frac{145665}{16384}\zeta[5] + \frac{20655}{8192}\zeta[3]\zeta[5] - \frac{496125}{131072}\zeta[7]\end{aligned}$$

(unproven) observation: only rationals and odd zetas

Summary of results

Complete trans-series expansion of the energy density:

$$\tilde{\varepsilon} = \sum_{n=0}^{\infty} f_n^{(+)}(v) \nu^{kn} \quad (k \text{ } N\text{-dependent})$$

$$f_0^{(+)} = f_0 \quad (\text{the original PT series})$$

NP expansion parameter: $\nu = e^{-2B}$

running coupling:

$$\frac{1}{v} + \gamma \ln v = 2B \quad \gamma = \frac{4-N}{N-2} \quad \nu = v^{-\gamma} e^{-\frac{1}{v}}$$

Using Volin's results we found an algorithm to determine all $f_n^{(+)}(v)$ algebraically

Main conjecture

$$\varepsilon_{\text{phys}} = \varepsilon_{\text{TBA}} = \mathcal{S}_+(\tilde{\varepsilon}) = \sum_{n=0}^{\infty} \nu^{kn} \mathcal{S}_+(f_n^{(+)})$$

no proof

but checked numerically to 72(56) digits for $N = 4(3)$

proved

$$\mathcal{S}_+(\tilde{\varepsilon}) = \mathcal{S}_{\text{med}}(f_0) \quad (N > 3)$$

$\varepsilon_{\text{phys}} \neq \mathcal{S}_+(f_0)$ but f_0 “knows” about all NP corrections

(f_0 determines all higher $f_n^{(+)}$)

$$N = 3 \ (k = 1)$$

f_0, f_1, f_2 not related

$$\varepsilon_{\text{phys}} = \varepsilon_{\text{TBA}} = \sum_{n=0}^{\infty} \nu^n \mathcal{S}_+(f_n^{(+)}) = \mathcal{S}_{\text{med}}(f_0 + \nu f_1 + \nu^2 f_2)$$

physical interpretation:

f_i : i instanton contribution $i = 0, 1, 2$

TBA(like) integral equation

TBA(-like) integral equation:

$$\chi_n(\theta) - \int_{-B}^B d\theta' K(\theta - \theta') \chi_n(\theta') = \cosh n\theta \quad |\theta| \leq B \quad n \geq 0$$

TBA context: 2-dim integrable QFT;

$K(\theta)$: logarithmic derivative of S-matrix

$(\mu \cosh \theta, \mu \sinh \theta)$ relativistic 2-momentum

B : (bosonic) fermi-rapidity

densities:

$$O_{n,m} = \frac{1}{2\pi} \int_{-B}^B d\theta \chi_n(\theta) \cosh m\theta \quad m \geq 0$$

explicit B -dependence dropped:

$$\chi_n(\theta) \sim \chi_n(\theta, B) \quad O_{n,m} \sim O_{n,m}(B)$$

Relativistic QFT:

$$\varepsilon = \mu^2 O_{1,1} \quad \rho = \mu O_{1,0} \quad \{\varepsilon(B); \rho(B)\} \implies \varepsilon = \varepsilon(\rho)$$

Differential relations among densities

$$O_{n,m} = O_{m,n} \quad (I) \quad \dot{O}_{n,m} = \frac{1}{\pi} \rho_n \rho_m \quad \dot{\phi} = \frac{d\phi}{dB}$$

where

$$\rho_n = \chi_n(B) \quad [\sim \chi_n(B, B)]$$

$$(II) \quad \ddot{\rho}_n - n^2 \rho_n = \mathcal{F} \rho_n \quad \mathcal{F} = \mathcal{F}(B) \quad \{n \text{ --independent}\}$$

Ristivojevic '22

1 density determines all:

$$O_{1,1} \longrightarrow \rho_1 \longrightarrow \mathcal{F} \longrightarrow \rho_n \longrightarrow O_{n,m}$$

Wiener-Hopf solution of TBA integral equations

Riemann-Hilbert problem

$$1 - \tilde{K}(\omega) = \frac{1}{G_+(\omega)G_-(\omega)} \quad G_-(\omega) = G_+(-\omega)$$

$$O(N) \text{ NLS model} \quad \{\Delta = 1/(N-2)\}$$

$$\tilde{K}(\omega) = e^{-\pi\Delta|\omega|} \frac{\cosh \frac{\pi}{2}(1-2\Delta)\omega}{\cosh \frac{\pi}{2}\omega} \quad \{\tilde{K}(\omega) = e^{-\pi|\omega|} \quad O(3)\}$$

$$\sigma(\omega) = \frac{G_-(\omega)}{G_+(\omega)} : \text{ cut and poles along positive imaginary axis}$$

poles and residues:

$$\sigma(iz \pm \epsilon) \approx \mp i \frac{S_\ell}{z - 2\ell\xi_o} \quad \ell = 1, 2, \dots$$

$$\xi_o = \begin{cases} \frac{N-2}{2} & N \text{ even} \\ N-2 & N \text{ odd} \end{cases} \quad [\xi_o = 1 \text{ for } O(3), O(4)]$$

discontinuity:

$$\sigma(iz + \epsilon) - \sigma(iz - \epsilon) = -2i\beta(z)$$

$\beta(z)$: meromorphic with poles at $z = 2\ell\xi_o$

expansion in terms of running coupling:

$$2B = \frac{1}{v} + \gamma \ln v + L \qquad \gamma = 2\Delta - 1$$

$$e^{-2Bvx} \beta(vx) = e^{-x} \mathcal{A}(x) \qquad \{ \mathcal{A}(x) \sim \mathcal{A}(x, v) \}$$

$$\mathcal{A}(x) = 1 + \sum_{k=1}^{\infty} (vx)^k L_k(\ln x)$$

(L_k polynomial of degree k)

$\nu = e^{-2B}$: NP expansion parameter v : PT expansion parameter

$$c_n = \nu^n \sigma(in + \epsilon)$$

reduced densities:

$$O_{n,m} = \frac{1}{4\pi} G_+(in) G_+(im) e^{(n+m)B} W_{n,m}$$

$$\rho_n = \frac{1}{2} G_+(in) e^{nB} w_n$$

WH integral equation:

$$Q_n(x) + \frac{c_n}{n+vx} + i \sum_{\ell=1}^{\infty} \frac{\nu^{2\ell\xi_o} S_{\ell} q_{n,\ell}}{2\ell\xi_o + vx} + \frac{1}{\pi} \int_{\mathcal{C}_+} \frac{e^{-y} \mathcal{A}(y) Q_n(y)}{x+y} dy = \frac{1}{n-vx}$$

$$q_{n,\ell} = Q_n \left(\frac{2\ell\xi_o}{v} \right)$$

$\mathcal{C}_+$: (just) above the real line

densities:

$$w_n = 1 + c_n + i \sum_{\ell=1}^{\infty} \nu^{2\ell\xi_o} S_{\ell} q_{n,\ell} + \frac{v}{\pi} \int_{\mathcal{C}_+} e^{-x} \mathcal{A}(x) Q_n(x) dx$$

conditions: $n, m > 0$ $n \neq m$ $n, m \neq 2\ell\xi_o$

$$W_{n,m} = \frac{1}{n+m} + \frac{c_m - c_n}{n-m} + c_m k_{n,m} + i \sum_{\ell=1}^{\infty} \frac{\nu^{2\ell\xi_o} S_{\ell} q_{n,\ell}}{m-2\ell\xi_o} + \frac{v}{\pi} \int_{\mathcal{C}_+} \frac{e^{-x} \mathcal{A}(x) Q_n(x)}{m-vx} dx$$

$$k_{n,m} = -\frac{c_n}{n+m} - i \sum_{\ell=1}^{\infty} \frac{\nu^{2\ell\xi_o} S_{\ell} q_{n,\ell}}{m+2\ell\xi_o} - \frac{v}{\pi} \int_{\mathcal{C}_+} \frac{e^{-x} \mathcal{A}(x) Q_n(x)}{m+vx} dx$$

Building blocks: TBA asymptotic series and transseries

Purely perturbative:

$$P_\alpha(x) + \frac{1}{\pi} \int_{\mathcal{C}_+} \frac{e^{-y} \mathcal{A}(y) P_\alpha(y)}{x+y} dy = \frac{1}{\alpha - vx}$$

well-defined (as asymptotic series) for all $\alpha \neq 0$

$$P_{\alpha,\beta} = \frac{v}{\pi} \int_{\mathcal{C}_+} \frac{e^{-x} \mathcal{A}(x) P_\alpha(x)}{\beta - vx} dx$$

well-defined (as asymptotic series) for all $\alpha, \beta \neq 0$
perturbatively:

$$W_{\alpha,\beta} \longrightarrow A_{\alpha,\beta} = \frac{1}{\alpha + \beta} + P_{\alpha,\beta}$$

well-defined for all $\alpha, \beta, \alpha + \beta \neq 0$

$$w_\alpha \longrightarrow a_\alpha = 1 + \frac{v}{\pi} \int_{\mathcal{C}_+} e^{-x} \mathcal{A}(x) P_\alpha(x) dx = \lim_{\beta \rightarrow \infty} \beta A_{\alpha,\beta}$$

differential relations for the reduced densities:

$$(n + m)W_{n,m} + \dot{W}_{n,m} = w_n w_m$$
$$\ddot{w}_n + 2n\dot{w}_n = \mathcal{F}w_n$$

perturbative limit:

$$(n + m)A_{n,m} + \dot{A}_{n,m} = a_n a_m$$
$$\ddot{a}_n + 2n\dot{a}_n = \mathcal{F}_o a_n$$

Volin's result determines all:

$$A_{1,1} \longrightarrow a_1 \longrightarrow \mathcal{F}_o \longrightarrow a_n \longrightarrow A_{n,m}$$

Volin's result for $A_{1,1}$:

$$2A_{1,1} = 1 + \frac{v}{2} + \left(\frac{5\gamma}{4} + \frac{9}{8}\right) v^2 + \left(\frac{10\gamma^2}{3} + \frac{53\gamma}{8} + \frac{57}{16}\right) v^3 \\ + \frac{v^4}{384} \left(-36\gamma^3(21\zeta_3 - 94) + 10924\gamma^2 + 13344\gamma + 9(144z_3 + 625)\right) + \dots$$

where

$$z_{2k+1} = 2 \frac{\zeta_{2k+1}}{2k+1} (\Delta^{2k+1} - 1 + 2^{-2k-1}) \quad .$$

next step

$$a_1 = 1 + \frac{v}{4} + \left(\frac{5\gamma}{8} + \frac{9}{32}\right) v^2 + \left(\frac{5\gamma^2}{3} + \frac{53\gamma}{32} + \frac{75}{128}\right) v^3 \\ + \frac{v^4(-288\gamma^3(21\zeta_3 - 94) + 43696\gamma^2 + 35160\gamma + 9(1152z_3 + 1225))}{6144} + \dots$$

next step

$$\mathcal{F}_o = -v^2 - 6\gamma v^3 - 26\gamma^2 v^4 + v^5 \left(\frac{1}{4}\gamma^3(63\zeta_3 - 386) - 27z_3\right) + \dots$$

next step

$$a_n = 1 + \frac{v}{4n} + \frac{v^2(20\gamma n+9)}{32n^2} + \frac{v^3(640\gamma^2 n^2+636\gamma n+225)}{384n^3} \\ + \frac{v^4(288n^3(\gamma^3(94-21\zeta_3)+36z_3)+43696\gamma^2 n^2+35160\gamma n+11025)}{6144n^4} + \dots$$

finally

$$A_{n,m} = \frac{1}{m+n} + \frac{v}{4mn} + \frac{v^2(20\gamma mn+9m+9n)}{32m^2 n^2} \\ + \frac{v^3(m^2(640\gamma^2 n^2+636\gamma n+225)+6mn(106\gamma n+39)+225n^2)}{384m^3 n^3} + \dots$$

trans-series solution of the linear problem:

$$\tilde{Q}_n(x) = P_n(x) + c_n P_{-n}(x) + i \sum_{\ell=1}^{\infty} \nu^{2\ell\xi_o} S_{\ell} \tilde{q}_{n,\ell} P_{-2\ell\xi_o}(x)$$

Complete (trans-series) solution:

$$\tilde{q}_{n,s} - i \sum_{\ell=1}^{\infty} \nu^{2\ell\xi_o} S_{\ell} \tilde{q}_{n,\ell} A_{-2\ell\xi_o, -2s\xi_o} = A_{n, -2s\xi_o} + c_n A_{-n, -2s\xi_o}$$

$$\begin{aligned} \tilde{W}_{n,m} = & A_{n,m} + i \sum_{\ell=1}^{\infty} \nu^{2\ell\xi_o} S_{\ell} \tilde{q}_{n,\ell} A_{-2\ell\xi_o, m} + c_n A_{-n, m} + c_m A_{n, -m} \\ & + c_n c_m A_{-n, -m} + i c_m \sum_{\ell=1}^{\infty} \nu^{2\ell\xi_o} S_{\ell} \tilde{q}_{n,\ell} A_{-2\ell\xi_o, -m} \end{aligned}$$

$A_{\alpha, \beta}$ (PT) and Stokes constants S_{ℓ} (NP) universal building blocks!

Energy density $\tilde{W}_{1,1}$

$$n \approx 1 \quad c_n = \nu \left\{ h_o + \frac{M}{2}(1 - n) + O((1 - n)^2) \right\} \quad h_0 = \frac{\delta_{N,3}}{e}$$

$N \geq 4$ and $N = 3$ cases drastically different!

$$M = \begin{cases} -2e \left(\frac{\Delta}{e}\right)^{2\Delta} \frac{\Gamma(1-\Delta)}{\Gamma(1+\Delta)} e^{i\pi\Delta} & N \geq 4 \\ -\frac{i\pi}{e} + \frac{2}{e}(2B + 1 - \gamma_E - \ln 2) & N = 3 \end{cases}$$

$$n = 1 \quad m = 1 + \delta \rightarrow 1$$

$$N \geq 4$$

$$\tilde{q}_{1,s} - i \sum_{\ell=1}^{\infty} \nu^{2\ell\xi_o} S_{\ell} \tilde{q}_{1,\ell} A_{-2\ell\xi_o, -2s\xi_o} = A_{1, -2s\xi_o}$$

$$\tilde{W}_{1,1} = A_{1,1} + \frac{M\nu}{2} + i \sum_{\ell=1}^{\infty} \nu^{2\ell\xi_o} S_{\ell} \tilde{q}_{n,\ell} A_{1, -2\ell\xi_o}$$

$$N = 3$$

$$\tilde{q}_{1,s} - i \sum_{\ell=1}^{\infty} \nu^{2\ell} S_{\ell} \tilde{q}_{1,\ell} A_{-2\ell, -2s} = A_{1, -2s} + \frac{\nu}{e} A_{-1, -2s}$$

$$\tilde{W}_{1,1} = A_{1,1} + \frac{M\nu}{2} + \frac{2\nu}{e} P_{1,-1} + \frac{\nu^2}{e^2} A_{-1, -1}$$

$$+ i \sum_{\ell=1}^{\infty} \nu^{2\ell} S_{\ell} \tilde{q}_{n,\ell} \left[A_{1, -2\ell} + \frac{\nu}{e} A_{-1, -2\ell} \right]$$

The main conjecture

$N > 3$ case

trans-series:

$$\tilde{\varepsilon} = 2\tilde{W}_{1,1} = 2A_{1,1} + M\nu + 2iS_1 A_{1,-k}^2 \nu^k + O(\nu^{2k}) \quad k = 2\xi_o$$

$$f_0 = 2A_{1,1}; \quad f_1^{(+)} = 2iS_1 A_{1,-k}^2; \quad f_2^{(+)} = 2iS_2 A_{1,-2k}^2 - 2S_1^2 A_{1,-k}^2 A_{-k,-k};$$

Main conjecture

$$\boxed{\varepsilon_{\text{phys}} = \varepsilon_{\text{TBA}} = \mathcal{S}_+(\tilde{\varepsilon})}$$

no proof but interesting consequences

need new definitions: Stokes automorphism, alien derivatives

Stokes automorphism:

$$\mathfrak{S} = \exp\{-\mathcal{R}\} \quad \mathcal{R} = \sum_{j=1}^{\infty} \nu^{kj} \Delta_{kj}$$

Δ_{kj} alien derivative at $\omega = kj$

The definition of alien derivatives are complicated in general, but for singularities closest to the origin in the Borel plane operationally simple:

$f(v) \longrightarrow B[f](t)$ has singularity (cut) at $t = \omega$

$B[\Delta_\omega f](t) \sim$ discontinuity across the cut

$$\mathcal{S}_-(X) = \mathcal{S}_+(\mathfrak{S}^{-1}X)$$

real X :

$$2i\mathrm{Im} \mathcal{S}_+(X) = \mathcal{S}_+[(1 - \mathfrak{S}^{-1})X]$$

Alternatively the alien derivative can be read off the the asymptotic behaviour of the perturbative coefficients:

$$\psi(z) = \sum_{n=0}^{\infty} \frac{\Gamma(n)c_{n-1}}{z^n} \quad \{z = 2B \text{ for } O(4)\}$$

$$B[\psi](t) = \sum_{n=0}^{\infty} c_n t^n \quad (\text{Borel transform})$$

If asymptotically:

$$c_n = \frac{A_o}{2} \left\{ 1 + \frac{p_0}{n} + \frac{p_1}{n(n-1)} + \frac{p_2}{n(n-1)(n-2)} + \dots \right\} \\ + \frac{B_o}{2} (-1)^n \left\{ 1 + \frac{q_0}{n} + \frac{q_1}{n(n-1)} + \frac{q_2}{n(n-1)(n-2)} + \dots \right\}$$

alien derivatives at 1 and -1 :

$$(\Delta_1 \psi)(z) = -i\pi A_o \left\{ 1 + \sum_{m=0}^{\infty} \frac{p_m}{z^{m+1}} \right\}$$

$$(\Delta_{-1} \psi)(z) = i\pi B_o \left\{ 1 - \sum_{m=0}^{\infty} \frac{(-1)^m q_m}{z^{m+1}} \right\}$$

reality of $\varepsilon_{\text{phys}}$:

$$\Delta_{kj} A_{1,1} = 2i S_j A_{1,-kj}^2$$

combining with differential relations:

$$\Delta_{kj} A_{n,m} = 2i S_j A_{n,-kj} A_{m,-kj}$$

complete resurgence structure: all alien derivatives of all building blocks are known

Main result

$$\tilde{\varepsilon} = M\nu + \mathfrak{S}^{-1/2} f_0$$

Median resummation:

$$\mathcal{S}_{\text{med}}(X) = \mathcal{S}_+(\mathfrak{S}^{-1/2} X) = \mathcal{S}_-(\mathfrak{S}^{1/2} X)$$

Aniceto, Schiappa '13

finally:

$$\varepsilon_{\text{phys}} = M\nu + \mathcal{S}_{\text{med}}(f_0)$$

f_0 determines full trans-series, “knows” about all higher NP contributions!

$$N = 3$$

trans-series:

$$\tilde{\varepsilon} = 2\tilde{W}_{1,1} = 2A_{1,1} + M\nu + \frac{4}{e}\nu P_{1,-1} + \frac{2}{e^2}\nu^2 A_{-1,-1} + \dots$$

Main conjecture $\longrightarrow$ same resurgence as $N \geq 4$ ($k = 2$)

using Stokes automorphism

$$\tilde{\varepsilon} = \mathfrak{S}^{-1/2}(f_0 + \nu f_1 + \nu^2 f_2)$$

$$f_0 = 2A_{1,1} \quad f_1 = M_o + \frac{4}{e}P_{1,-1} \quad f_2 = \frac{2}{e^2}A_{-1,-1}$$

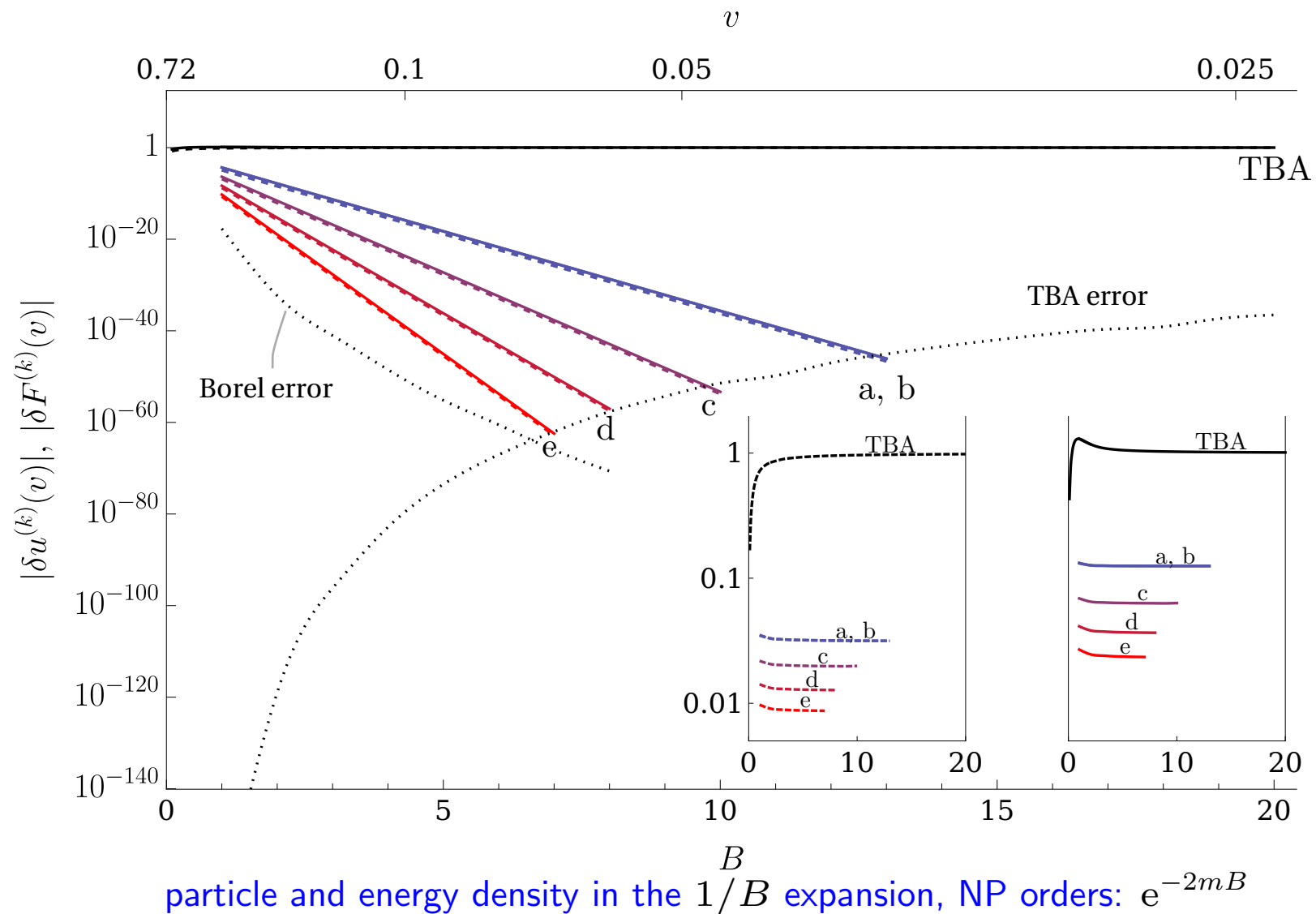
$$M_o = \frac{2}{e}(1 - \gamma_E - \ln 2 + 2B) = \frac{2}{e}\left(\frac{1}{v} + \ln v - \gamma_E - 4 \ln 2\right)$$

median resummation:

$$\varepsilon_{\text{phys}} = \mathcal{S}_{\text{med}}(f_0 + \nu f_1 + \nu^2 f_2)$$

need f_0 (PT) and 1 and 2-instanton sectors

Numerical results



conclusion:

$$f_{\text{exact}} = f_{\text{TBA}} = S_{\text{med}}(f)$$

Convergence? YES

Thank you!